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Technical Report No. 47

Utilization of a Highly Correlated Cluster Model for
Interpretation of Electronic Spectroscopic Data for
The High-Temperature Superconductors

By

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Prepared for Publication

in the

AIP/AVS Topical Conference Proceedings

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December, 1988

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SECURITY CLASSIFICATION OF THIS PAGE

REPORT DOCUMENTATION PAGE

1a. REPORT SECURITY CLASSIFICATION Unclassified		1b. RESTRICTIVE MARKINGS		
2a. SECURITY CLASSIFICATION AUTHORITY		3. DISTRIBUTION/AVAILABILITY OF REPORT Approved for Public Release, distribution Unlimited.		
3b. DECLASSIFICATION/DOWNGRADING SCHEDULE				
4. PERFORMING ORGANIZATION REPORT NUMBER(S) Technical Report # 47		5. MONITORING ORGANIZATION REPORT NUMBER(S)		
6a. NAME OF PERFORMING ORGANIZATION Dept. of Chemistry George Washington Univ.	6b. OFFICE SYMBOL (If applicable)	7a. NAME OF MONITORING ORGANIZATION Office of Naval Research (Code 413)		
6c. ADDRESS (City, State, and ZIP Code) Washington, D.C. 20052		7b. ADDRESS (City, State, and ZIP Code) Chemistry Program 800 N. Quincy Street Arlington, VA 22217		
8a. NAME OF FUNDING/Sponsoring ORGANIZATION Office of Naval Research	8b. OFFICE SYMBOL (If applicable)	9. PROCUREMENT INSTRUMENT IDENTIFICATION NUMBER Contract N00014-80-K-0852		
8c. ADDRESS (City, State, and ZIP Code) Chemistry Program 800 North QUINCY, Arlington, VA 22217		10. SOURCE OF FUNDING NUMBERS		
		PROGRAM ELEMENT NO. 61153 N	PROJECT NO. TASK NO. PP 013-08-01	ACCESSION NO. NR 056-681
11. TITLE (Include Security Classification) Utilization of a Highly Correlated Cluster Model for Interpretation of Electronic Spectroscopic Data for the High-Temperature Superconductors. (Unclassified)				
12. PERSONAL AUTHOR(S) D. E. Ramaker				
13a. TYPE OF REPORT Interim Technical	13b. TIME COVERED FROM _____ TO _____	14. DATE OF REPORT (Year, Month, Day) December 1988	15. PAGE COUNT 7	
16. SUPPLEMENTARY NOTATION Prepared for publication in AIP/AVS Topical Conference Proceedings				
17. COSATI CODES REF ID: GROUP SUB-GROUP	18. SUBJECT TERMS (Continue on reverse if necessary, and identify by block number) Superconductivity, Hubbard Model, Photoelectron Spectroscopy, Auger Spectroscopy, Copper Oxides, (Many)			
19. ABSTRACT (Continue on reverse if necessary and identify by block number) We have consistently interpreted electron spectroscopic data for the high temperature superconductors utilizing a highly-correlated CuO _n ⁷ cluster model, and an extended Hubbard Hamiltonian which includes the inter-site Cu-O and O-O U _p ⁷ parameters. The data indicate much larger U _p ⁷ and U _{pp} ⁷ values than found in other typical highly conductive metals. Previously unassigned features in the data are now assigned within the model.				
20. DISTRIBUTION/AVAILABILITY OF ABSTRACT <input checked="" type="checkbox"/> UNCLASSIFIED/UNLIMITED <input checked="" type="checkbox"/> SAME AS RPT. <input type="checkbox"/> DDCR USERS		21. ABSTRACT SECURITY CLASSIFICATION Unclassified		
22a. NAME OF RESPONSIBLE INDIVIDUAL Dr. David L. Nelson		22b. TELEPHONE (Include Area Code) (202) 696-4410	22c. OFFICE SYMBOL	

DD FORM 1473, 24 MAR

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UTILIZATION OF A HIGHLY-CORRELATED CuO_n CLUSTER MODEL TO
INTERPRET ELECTRON SPECTROSCOPIC DATA FOR THE HIGH-
TEMPERATURE SUPERCONDUCTORS

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ABSTRACT

We have consistently interpreted electron spectroscopic data for the high temperature superconductors utilizing a highly-correlated CuO_n cluster model, and an extended Hubbard Hamiltonian which includes the inter-site Cu-O U_{pd} and O-O U_{oo} parameters. The data indicate much larger U_{pd} and U_{oo} values than found in other typical highly conductive metals. Previously unassigned features in the data are now assigned within the model.

INTRODUCTION

In this work we summarize results of an interpretation of electron spectroscopic data for the high temperature superconductors. The data interpreted include the valence band (VB), Cu 2p, and O 1s photoelectron data (UPS and XPS), the Cu L_{2,3}VV, Cu L_{2,3}M_{2,3}V, and O KVV Auger data, and the O K and Cu L_{2,3} x-ray absorption and emission (XANES and XES) data. Published data for polycrystalline and single crystal samples of La_{2-x}Ba_xCuO₄ and YBa₂Cu₃O_{7-x} (herein referred to as the La and 123 HTSC's) are considered along with that for CuO and Cu₂O.

The basic electronic structure of the HTSC's can be described with the Anderson Hamiltonian utilized by Sawatzky and coworkers^{1,2}. It includes the transfer or hopping integral t, the Cu and O orbital energies c_d and c_p, the core polarization energy Q_{dd}, and the intra-site Coulomb repulsion energies U_{dd} and U_{pp} (the latter sometimes are assumed to be zero). This model is most useful when the U's are large relative to the band widths¹, i.e. when correlation effects dominate hybridization effects. A CuO_n⁽ⁿ⁼⁴⁾ cluster model, which is also reasonably valid when U >> t, simplifies the model further¹. We utilize an extended Hubbard model by adding the inter-site repulsion energies U_{pd} and U_{oo}* (i.e. between neighboring Cu-O and O-O atoms). The addition of these interactions is important for understanding many of the features in the data.

RESULTS

Our results for the Hubbard parameters are summarized at the top of Table I. Other estimates of these Hubbard U and c parameters have been reported previously for the HTSC's²⁻⁴. These were obtained empirically from the Cu 2p XPS and the VB UPS data utilizing the Anderson model. Our optimal extended Hubbard

*Supported in part by the Office of Naval Research.

results indicate that $U_0 = 12$ and $U_{pp^*} = 4.5$ eV for 123. These are much larger than previously thought for these metallic systems, although $U_0 - U_{pp^*}$ is in agreement with the best theoretical results above.

An upper estimate of the two-center pp^* hole-hole repulsion, U_{pp^*} , can be obtained from the Klopman approximation⁷,

$$U_0 = e^2/(ru^2 + (2e^2/(U_0 + U_1))^2)^{1/2}, \quad (1)$$

where ru is the interatomic distance and U_0 and U_1 are the corresponding intra-atomic repulsion energies. Equation 1 gives a value for U_{pp^*} around 4.8 eV assuming r_{e-e} is 2.7 Å. The experimental energies of 9.5 and 5.0 eV for the pp^* and pp^* features in 123 suggests that the pp^* final state energy is 7.2 eV. This gives an empirical estimate for U_{pp^*} of 4.2 eV, very close to the Klopman theoretical result, which does not include the effects of interatomic screening.

The above result shows that metallic screening of two holes, which are spatially separated on neighboring O atoms, is not very significant. This is in contrast to two Cu-O holes, where Table 1 indicates the optimal $U_{pp^*} = 1$ eV, while eq. 1 estimates U_{pp^*} at 6.1 eV assuming r_{e-e} is 1.9 Å. This large reduction in U_{pp^*} may result from charge transfer into the Cu 4sp levels to screen the Cu-O holes. Although metallic screening, which results from virtual electron-hole (e-p) pair excitations at the Fermi level, is not expected to be large in an insulator such as CuO, screening effects are expected to be much larger in metals, such as the HTSC's. The above results show that U_{pp^*} is significantly reduced in both, and U_{pp^*} remains large in both. The lack of a significant change in the U 's between CuO and the HTSC's indicates that the DOS at the Fermi level in the HTSC's must be very small.

Table 1 also correlates the calculated energies of the excited states with features in the experimental data. The $CuO_x^{(2+2)}$ cluster has one hole shared between the Cu 3d and O 2p shells in the ground state, which we term the v (valence) states. The spectroscopic final states reflect multi-hole states, e.g. v^2 , $c v$ (c = core) etc. The v states, as reflected by the theoretical DOS⁸, have the Cu-O bonding (+) and antibonding (+) orbitals centered at 4 and 0 eV, respectively, and the nonbonding Cu and O orbitals at 2 eV. The O features each have a width $2\Gamma = 4$ eV due to the O-O bonding and antibonding character and the Cu-O dispersion. We also define the Cu-O hybridization shift $\delta_1 = ((\Delta + 4t^2)/2 - \Delta)/2$, which is utilized in Table 1 to give the energies. Thus, the ground state of an average CuO_x cluster is located at 1 eV having the energy $\epsilon_0 - \delta_1 + \Gamma/2 = c_4 - \alpha$, which we use as a reference energy for the excited states. In CuO, the hybridization shift Γ is smaller, and we shall see below that $\Delta = c_0 - c_4$ has increased to 1 eV.

Those clusters containing additional charge carrier holes (these exists in doped La, and 123 when $x > 0.5$) actually have two holes per CuO_x cluster. The average v^2 ground state, which is dominated by the pp^* configuration, i.e. the charge carrier holes spend most of their time on the O atoms, so we indicate this ground state by the

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TABLE I Summary of hole states revealed in the spectroscopic data, and estimated energies using the following optimal values for the Hubbard parameters in eV:

$\delta_1 = 2$ $\epsilon_d = 2$ $U_p = 12, 13$ $U_d = 9.5, 10.2$
 $\delta_1 = 0.5, 0.8$ $\epsilon_p = 2, 3$ $U_{pp} = 4.5, 4$ $U_{dp} = 1$
 $\Gamma = 2$ $U_{pd} = 0$ $U_{cp} = 2$ $Q_d = 9$
 $\alpha = 1, 0.5$ $\beta = 2$ $\Delta = 0, 1.$ $K = 4$

State*	Energy expression eV ^{c,d}	Calc. E. eV ^c	Exp. E. eV ^c	Remark
<u>G.S. and IPES, v²</u>				
+ _a) d	$\epsilon_d - \delta_1 \mp \Gamma$	0 \mp 2	-	heavily mixed
+ _b) p	$\epsilon_p + \delta_1 \mp \Gamma$	4 \mp 2	-	
<u>UPS and XES, v²</u>				
1) pp ^p	$\epsilon_p + \Delta - \delta_2 + \alpha$	2.5	2.5	heavily mixed
2) dp	$\epsilon_p + U_{pp} + \delta_2 + \alpha$	4.5	4.2	
3) pp ^{p*}	$\epsilon_p + \Delta + U_{pp} - \Gamma + \alpha$	5.5	5.	
4) pp ^{p*} *	$\epsilon_p + \Delta + U_{pp} + \Gamma + \alpha$	9.5	9.5	mystery peak
5) d ²	$\epsilon_d + U_d + \alpha$	12.5	12.5	Cu sat.
6) p ²	$\epsilon_p + \Delta + U_p + \alpha$	15	16	
<u>Cu 2p XPS, cv</u>				
d \rightarrow cp	$\epsilon_c + \Delta + \alpha$	$\epsilon_c + 1$	E_{1s}	main
cd	$\epsilon_c + Q_d + \alpha$	$\epsilon_c + 10$	$E_{1s} + 9.2$	sat.
<u>Cu 2p XPS for NaCuO₂, pp^p \rightarrow cv²</u>				
pp ^p \rightarrow cpp ^p	$\epsilon_c + \delta_2 + \beta$	$\epsilon_c + 2.5$	$\epsilon_c + 2.2$	main
	$\epsilon_c + U_{pp} - \Gamma + \delta_2 + \beta$	$\epsilon_c + 4.5$	$\epsilon_c + 5$?
	$\epsilon_c + U_{pp} + \Gamma + \delta_2 + \beta$	$\epsilon_c + 8.5$	$\epsilon_c + 9$?
	$\epsilon_c - \Delta + Q_d + U_{pp} + \delta_2 + \beta$	$\epsilon_c + 11.5$	$\epsilon_c + 11$	sat.
	$\epsilon_c + U_p + \delta_2 + \beta$	$\epsilon_c + 15.5$	$\epsilon_c + 14$	sat.?
<u>O 1s XPS, cv</u>				
d \rightarrow cd	$\epsilon_c + \alpha$	$\epsilon_c + 1$	E_{1s}	main
cp ^p	$\epsilon_c + \Delta + \alpha$	$\epsilon_c + 1$	E_{1s}	main
cp ^{p*}	$\epsilon_c + \Delta + U_{cp} + \alpha$	$\epsilon_c + 3$	$E_{1s} + 2$?	tail
cp	$\epsilon_c + \Delta + Q_p + \alpha$?	?	not obs
pp ^p \rightarrow cdp ^p	$\epsilon_c - \Delta + U_{cp} + \delta_2 + \beta$	$\epsilon_c + 3.5$	$E_{1s} + 2$?	tail
<u>Cu L₃VV ABS, v²</u>				
dpp ^p	$2\epsilon_p + 2U_{dp} + \alpha$	7	7	2 cent.
dpp ^{p*}	$2\epsilon_p + U_{pp} + 2U_{dp} + \alpha$	11.5	-	no mix
d ² p	$\epsilon_d + \epsilon_p + U_d + 2U_{dp} - \delta_2 + \alpha$	16	15.5	main
dp ^p	$2\epsilon_p + U_p + 2U_{dp} + \delta_2 + \alpha$	19.5	18-25	sat.

TABLE 1 (cont.)

State ^b	Energy expression	Calc. E. eV ^c	Exp. E. eV ^c	Remark
<u>Cu L_{2,3}V AES, cv^d</u>				
cdp	$\epsilon_c + \epsilon_p + Q_d + U_{dp} + K + \alpha$	$\epsilon_c + 9$ $\epsilon_c + 17$	$E_{2p} + 10$ $E_{2p} + 18$	main, ^f L
cp ^e	$\epsilon_c + \epsilon_p + \Delta + U_p + \alpha$	$\epsilon_c + 15$	-	not obs.
cd ^e	$\epsilon_c - \epsilon_p + U_d + 2Q_d + \alpha$	$\epsilon_c + 30.5$	-	obs.
<u>Cu L_{2,3} EELS, c</u>				
d → c	$\epsilon_c - \epsilon_d + \delta_1$	$E_{2p} - 1$	$E_{2p} - 1.4$	edge
cpCB	$\epsilon_c + \Delta - CB + \alpha$	$E_{2p} - CB$	$E_{2p} + 1.2$	upper
pp ^f → cp	$\epsilon_c - \epsilon_p + \delta_1 + \delta$	$E_{2p} - 0.5$	E_{2p}	middle
<u>O K EELS, c</u>				
d → c	$\epsilon_c - \epsilon_d + \delta_1$	$E_{1s} - 1$	E_{1s}	edge
cdCB	$\epsilon_c - CB + \alpha$	$E_{1s} - CB$	$E_{1s} + 1.7$	upper
pp ^f → cd	$\epsilon_c - \Delta - \epsilon_p + \delta_1 + \delta$	$E_{1s} - 0.5$	-	not obs.

^aParameters for 123 indicated first, those for CuO second.^bThe dominant character in the hybridized states is given.^cThe Calc. E and Exp. E columns indicate the results for 123, except for the "Cu 2p XPS, pp^f → cv^d" section, which is for NaCuO₂.^dThe calculated E is defined relative to the ground vⁱ (d) state energy = $\epsilon_d - \alpha$, or to the vⁱ (pp) ground state energy = $2\epsilon_p - \delta - \alpha$. The vⁱ(d) energy defines the Fermi level relative to the vacuum level at zero.^eThe dominant character switches as described in the text, and thus the sign in front of δ_1 is the opposite for CuO.

parameters in Table 1 were obtained by considering this same data plus XANES, Auger and XES data. Although we are in general agreement with the reported magnitudes for most of the parameters, our U_d value is larger by about 2-3 eV so that it is consistent with the AES data. In Table 1, we indicate the location of two valence holes by d (Cu 3d) or p (O 2p). In the case of two holes on the oxygens, we distinguish two holes on the same O (p²), on ortho neighboring O atoms (pp²), or on para O atoms (pp⁴) of the cluster. Furthermore, neighboring pp⁴ holes can dimerize⁶, so we distinguish between two holes in bonded (pp⁴) and antibonded (pp⁴) O pairs.

The magnitudes of the U parameters are critical to the mechanism for the superconductivity. As a consequence, much effort has also gone into theoretically calculating these parameters, but wide disagreement still exists over the magnitudes. Theoretical values for U_d in the range 6.5-10 eV, U_p (actually $U_p - U_{pp^2}$) in the range 7-14 eV, and U_{dp} in the range 0.6-1.6 eV have been reported⁶, with the smaller results favored based on the quality of the calculations. No results for U_{pp^4} have been reported. Our empirical

notation pp^* . We use $2c_0 - d_0 - \delta$ as the energy of the pp^* ground state relative to the vacuum level, where $\delta = 2$ eV is the energy shift between the principal pp^* UPS final state at 2.5 eV and the lowest ground pp^* states around 0.5 eV from the Fermi level.

The correlation between the calculated energies and experimental features, utilizing the indicated optimal Hubbard parameters is very good. Details of this work are published elsewhere⁸. Figs. 1 and 2 show examples of the UPS and Cu AES data for 123 and CuO₂, which reveal some of the features itemized in Table 1; the remaining data are published elsewhere.

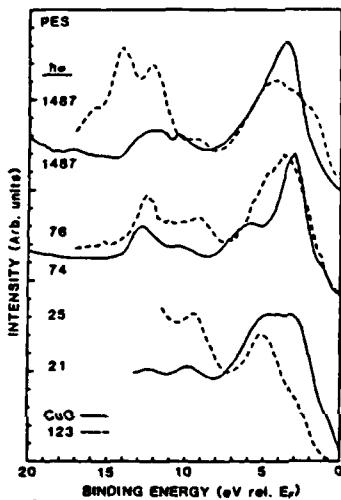


Figure 1. Comparison of UPS spectra for CuO and 123 taken with the indicated photon energies in eV. Data for CuO from refs. 10 ($h\nu = 1487$), 11 ($h\nu = 74$) and 12 ($h\nu = 21$). Data for 123 from ref. 13 ($h\nu = 25$ and 74) and 14 ($h\nu = 1487$).

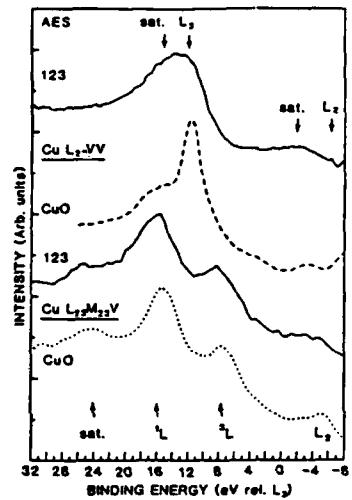


Figure 2. Comparison of Auger data for the materials indicated. L₂VV data for CuO and 123 from ref. 15. L₂M₂V data for CuO from ref. 16 and for 123 from Ref. 9. The L₂VV data is on a 2-hole binding energy scale = E_F-E_{L2}, and the L₂M₂V on a 1-hole scale = E_F-E_{L2}-E_{M2}, where E_{L2} = 933.4 and E_{M2} = 77.3 eV^{11,12}.

SUMMARY AND CONCLUSIONS

This work has allowed us to assign some previously unassigned features in the data, and greatly increased our understanding of the

dynamical electronic processes which produce these features. We itemize our conclusions as follows:

- 1) A switch in the character of state 1 (see Table 1) from more dp to pp^{*} and vice versa for state 2 between CuO and 123 arises because Δ decreases from 1 eV to 0 eV. The smaller Δ in 123, due to a smaller c_c , polarization energy, is consistent with the Cu 2p XPS and XES data (the latter showing this effect dramatically)^{17,18}. Since state 1 is more of pp^{*} character in the SC's, the "charge carrier holes" (present in the La after Sr doping and in the 123 when T_c is greater than 6.5) spend more time on the oxygens in 123 than in CuO.
- 2) The pp^{*} state is believed to be responsible for the "mystery" peak found at 9.5 eV in the UPS. Figure 1 indicates that such a feature also appears for CuO^{11,12}. This feature does not appear for Cu₂O, as expected since UPS reflects the one-hole DOS in Cu₂O. Thus this feature is not unique to the SC's; it naturally appears for those systems with two-hole photoemission final states.
- 3) Although Cu₂O, CuO, and NaCuO₂ have a formal Cu valence of +1, +2, and +3, in the current picture they reflect the cv_n DOS, with n=0, 1, and 2. Furthermore, we consistently predict the "chemical shifts" in the primary Cu 2p XPS peaks. Whereas, Cu₂O exhibits just a primary core hole c state at energy c_c , CuO has its primary cd feature energy shifted by $\Delta+\delta$ relative to c_c , and the primary cpp^{*} feature for NaCuO₂ by $\delta+\Delta$ (Table 1), which is consistent with the experimental data¹⁷. The width of the primary feature is seen to correlate with the intensity of the satellite, and is not due to the O p band width as suggested by others¹⁹.
- 4) The increased "satellite" feature at 19 eV in the Cu L_{2,3}VV Auger line shape for the HTSC's compared with CuO^{13,14} (see Fig. 2) arises because of increased final-state configuration mixing between the d³p and d²p states. Its intensity is increased in 123 relative to CuO because the energy separation (before hybridization) between d³p and d²p has decreased from 3.8 eV in CuO to 2.5 eV in 123. We have indicated this mixing in Table 1 by adding the hybridization shifts δ_1 to the energy expressions for these two states.
- 5) We find that the initial-core shakeup (ICSU) process, which is known to be responsible for the satellite features in the Cu 2p XPS data¹, does not produce satellites in the AES or XES data, because the ICSU states generally "relax" to the primary states of the same symmetry before the core level decay. Such a relaxation is expected when the ICSU excitation energy is larger than the core level width¹⁹. Previously, vanderLaan et al.¹, for the Cu halides, suggested that the intensity of these ICSU states in the XPS should be quantitatively reflected in the intensity of the Auger satellites found in the L_{2,3}VV lineshapes. The data do not indicate this however. We previously¹ indicated that a fraction of these ICSU states probably resulted in Auger satellites for the HTSC's, and that this fraction becomes larger as the covalency of the HTSC material increases. This work indicates rather that the ICSU states relax before the core level decay to states of the same symmetry, provided they have a ICSU excitation energy that is much greater

than the core level width. We believe this to be a general result, at least in the Cu²⁺ materials.

6) The EELS and XANES data²⁰⁻²¹ reflect the contributions from three possible transitions; the dominant d → c contribution nearest the Fermi level, the pp^o → cv (v = d or p) contribution resulting from the carrier hole states, as well as the cvCB contribution well above the Fermi level²². Here CB represents an electron present in the higher Cu 4sp or O 3p "conduction band". The latter two contributions are not always resolved, and sometimes have been confused in the literature²⁰⁻²¹.

7) All of the temperature effects seen in the spectroscopic data²²⁻²⁴ can be attributed to a single phenomenon, namely a decrease in ϵ_p , due to increased metallic screening, or long range polarization. This is consistent with the decrease in the primary cp peak energy in the Cu 2p XPS, while the cd satellite remains unshifted. The larger energy separation between the cd and cp states decreases the mixing which causes the satellite to decrease in intensity and the main peak to get narrower. Although the primary cd peak does not shift in the O 1s XPS, a slight shift to lower energy is seen in the cp^o and cp^e contributions at lower temperature, as expected with a decrease in ϵ_p . The UPS spectra show a skewing toward the Fermi level at lower temperature, as expected with a decrease in ϵ_p . Finally the growth of the satellite intensity in the Cu L_{2,3}VV Auger lineshape is consistent with a decrease in ϵ_p . The increased metallic like screening or polarization which appears to occur at lower temperature, reducing ϵ_p , probably involves the grain boundaries, since the more recent data for the single crystal samples do not change with temperature²⁴.

In summary, an interpretation of the data utilizing a highly correlated CuO₆ cluster model shows that a single set of Hubbard parameters predicts all of the state energies. Changes in the data between CuO and the HTSC's arises primarily from a reduction in ϵ_p ; this reduction continues with decreasing temperature in the HTSC's due to increased metallic screening. Compared with CuO, the HTSC's show an increased covalent interaction between the Cu-O bonds. The large size of U_{pp}, and the temperature dependence, reveal that metallic screening is incomplete, and hence that the DOS at the Fermi level in the HTSC's is relatively small.

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